## [tex23] Thermodynamics of black-body radiation

Electromagnetic radiation inside a cavity is in thermal equilibrium with the walls at temperature T. The radiation has an energy density that depends only on the temperature, i.e. its internal energy has the form U(T, V) = Ve(T). The radiation pressure is determined by the energy density alone:  $p = \frac{1}{3}e(T)$ . (a) Use the consistency equations for the total differential dS to show that the energy density has the form  $e(T) = \sigma T^4$ , where  $\sigma$  is a constant, now known as the Stefan-Boltzmann constant. In this argument, the additional assumption enters that  $e(T) \to 0$  for  $T \to 0$ . (b) Determine the entropy S(T, V) and the thermodynamic potentials U(S, V), E(S, p), A(T, V), G(T, p). (c) Determine the isotherms and adiabates in the (V, p)-plane. (d) Determine the response functions  $C_V, C_p, \kappa_T, \kappa_S, \alpha_p$ .

## Solution: