

## [tex23] Thermodynamics of black-body radiation

Electromagnetic radiation inside a cavity is in thermal equilibrium with the walls at temperature  $T$ . The radiation has an energy density that depends only on the temperature, i.e. its internal energy has the form  $U(T, V) = Ve(T)$ . The radiation pressure is determined by the energy density alone:  $p = \frac{1}{3}e(T)$ .

- (a) Use the consistency equations for the total differential  $dS$  to show that the energy density has the form  $e(T) = \sigma T^4$ , where  $\sigma$  is a constant, now known as the *Stefan-Boltzmann constant*. In this argument, the additional assumption enters that  $e(T) \rightarrow 0$  for  $T \rightarrow 0$ . Infer the function  $S(T, V)$ .
- (b) Infer the thermodynamic potentials  $U(S, V)$ ,  $E(S, p)$ ,  $A(T, V)$ ,  $G(T, p)$ .
- (c) Determine the isotherms and adiabates in the  $(V, p)$ -plane.
- (d) Determine the response functions  $C_V, C_p, \kappa_T, \kappa_S$ .

**Solution:**