The Solubility of Spin-Polarized ³He in ⁴He and the Superfluidity of ³He in ³He-⁴He Mixtures

Lars P. Roobol,* Giorgio Frossati, Kevin S. Bedell,† and Alexander E. Meyerovich‡

Kamerlingh Onnes Laboratorium, Rijksuniversiteit Leiden, P.O. Box 9506, NL-2300 RA, Leiden, The Netherlands

(Received September 26, 1994; revised March 20, 1995)

We investigate the possibility of a large enhancement of the T = 0 finite solubility of ³He in ⁴He due to spin-polarization. The size of the effect depends on the fraction of ³He atoms in the system. We present two different approaches for the limits of a small and a large number of ³He atoms compared to the number of ⁴He atoms. Since the possible ³He superfluid phase transition depends on ³He density, we calculate the consequences of this change in the solubility for its superfluid transition temperature. It is shown that for small fractions of ³He, the transition temperature is enhanced mostly due to the enlargement of the up-spin Fermi sphere. In the opposite limit the transition temperature is enhanced as a result of the increased ³He solubility. PACS: 67.60.Fp, 64.75. + g, 67.75. + z

1. INTRODUCTION

A superfluid transition of the ³He sub-system in ³He-⁴He mixtures remains one of the most intriguing problems of low-temperature physics, as it would provide us with a unique mixture of two distinct superfluids. At present, not only the transition temperature and its dependence on thermodynamic variables like pressure are unknown, but even the exact type of pairing and the symmetry of the order parameter are unclear. As a result, theoretical estimates for T_C differ by several orders of magnitude. Nevertheless, all theories agree on the point that the properties of this mixture of superfluids should be very different from those of pure ³He or ⁴He. This can justify the considerable experimental effort necessary to try to observe the ³He superfluid transition in ³He-⁴He mixtures.

^{*}Now at: Royal Holloway, University of London, Egham, Surrey, TW20 0EX, U.K. †Los Alamos National laboratory - T 11, MS-B262, Los Alamos, NM 87545, U.S.A. ‡Department of Physics, University of Rhode Island - Kingston, RI 02881, U.S.A.

Up until now, some laboratories have cooled mixtures of various concentrations down to temperatures of 0.1-0.2 mK,¹⁻⁵ but none of them succeeded in observing the ³He superfluid transition in mixtures. Because of a large number of experimental parameters (temperature *T*, pressure *P*, ³He concentration *X* and magnetic field *B*) it is necessary to investigate in which region of this parameter space one is most likely to observe this superfluid transition.

We argue that one of the best strategies, from both the theoretical and experimental point of view, is to study ${}^{3}\text{He}{}^{4}\text{He}$ mixtures with a relatively high degree of spin-polarization of the ${}^{3}\text{He}$ component. Recent achievements in the methods of spin polarization⁶⁻¹⁰ can make such an approach feasible.

As we will see, spin polarization can lead to considerable changes in either the limiting solubility of ³He in ⁴He or the effective (³He-) interaction, depending on the model chosen. Since the transition temperature depends on these quantities exponentially, the effect of spin polarization on the transition temperature can be rather dramatic.

As examples, we will consider the relevant effects in the frames of both the s-wave and the potential models. In some sense, these models are opposite to each other: the s-wave model uses an unrenormalized interaction and relates all the polarization changes to the change in density of states, while the main effect for the potential models is the change in effective interaction. Therefore, the s-wave model predicts large changes in solubility with polarization which causes the transition temperature to increase exponentially. On the other hand, the solubility changes for the potential models are less significant than the exponential dependence of the transition temperature on renormalized interaction. Despite these major differences between the models, both predict that spin polarization could cause a very significant increase in transition temperature. The fact that both these opposite models predict a large increase in T_c signals that the study of spin-polarized mixtures can be very promising.

As an example of the behaviour of a potential model, we will use the one originally proposed by Bardeen, Baym and Pines,¹¹ extended to finite polarization by van de Haar, Bedell and Frossati¹² for calculating the properties of the mixture. The corresponding properties of the pure phase are calculated using the "nearly metamagnetic model" put forward by Bedell and Sanchez–Castro¹³ using the (³He-) fit parameters of Sanchez–Castro, Bedell and Wiegers.¹⁴ We were able to solve the equations for phase equilibrium in the case where the volume of the mixture is much larger than that of the pure phase. In the opposite case, we were able to solve them using the dilute gas model.

In the next section, we briefly review the results on finite solubility in

zero magnetic field, extending it to finite field in Sec. III, using several different approaches. In Sec. IV we discuss the consequences for the transition temperature. It is shown that a possibly large enhancement of the solubility of spin-polarized ³He in ⁴He and the corresponding increase in density of states has a much larger effect on T_C than the increase in pairing interaction.

2. SOLUBILITY AT ZERO MAGNETIC FIELD

Below 0.8 K, one cannot mix ³He and ⁴He in arbitrary proportions: above a certain limiting concentration $x_s(T)$ of ³He, the mixture decomposes, as first observed by Walters and Fairbank,¹⁵ into dilute ³He-⁴He mixtures (dilute phase) and practically pure ³He (concentrated phase).

While the ³He-rich mixtures are self purifying (their ⁴He solubility limit becomes vanishingly small upon approaching zero temperature^{16,17}), the dilute mixtures have a finite solubility¹⁸ even at T = 0.

In this paper, we consider a fixed volume V containing a phase separated mixture at T=0 consisting of N_4 ⁴He atoms with N_{34} ³He atoms dissolved, and N_3 atoms in the concentrated phase. We define the total ³He concentration X and the ³He concentration $x \leq X$ in the mixture as

$$X = \frac{N_3 + N_{34}}{N_3 + N_{34} + N_4} \tag{1a}$$

$$x = \frac{N_{34}}{N_{34} + N_4} \tag{1b}$$

2.1. The Chemical Potentials

In thermodynamic equilibrium, the ³He chemical potentials of both phases must be equal, i.e.,

$$\mu_3(n_3, P_3) = \mu_{34}(n_{34}, P_{34}, n_4) \tag{2}$$

where n_3 (n_{34}) is the ³He density in the pure (dilute) phase, n_4 is the ⁴He density in the mixture and *P* is the pressure. Letting the pressure on both sides be equal fixes the density $n_s(P)$ at which Eq. (2) is satisfied.

In a fixed volume V, the chemical potentials are linked through the Gibbs–Duhem equations

$$dP_3 = n_3 \, d\mu_3 \tag{3a}$$

$$dP_{34} = n_{34} \, d\mu_{34} + n_4 \, d\mu_4 \tag{3b}$$

L. P. Roobol et al.

The last equation can be simplified for low ³He concentration using $d\mu_4 \approx dP/n_4^0$, with n_4^0 the density of *pure* ⁴He. The ³He chemical potential of the mixture contains three terms,

$$\mu_{34} = \varepsilon_0(\mu_4) + \frac{p_F^2}{2m^*(\mu_4)} + \varepsilon_{\rm int}(\mu_4, n_{34}) \tag{4}$$

where ε_0 is a constant depending on the ⁴He chemical potential μ_4 (or pressure), the second term is the Fermi energy of a gas of free particles with mass m^* , and Fermi momentum $p_F = \hbar k_F = \hbar (3\pi^2 n_{34})^{1/3}$. Finally, ε_{int} is a term due to the interaction between the ³He particles. This interaction term can be approximated by various methods. We will compare the results of three of such methods:

1. A fit to experimental data of Seligman *et al.*, 19,20 compiled in Ref. 21. Here it was assumed that the chemical potential can be expanded in the ³He density as

$$\mu_{34} = c_0 + c_1 n_{34}^{2/3} + c_2 n_{34} + c_3 n_{34}^{5/3} + \cdots$$
(5)

Since this approach also fits the two lowest order terms of expansion (4) in n_{34} , we do not have to use assumptions for the values of ε_0 and m^* .

2. The s-wave approximation, a microscopic approach in which the energy spectrum is calculated using perturbation theory. At all concentrations up to the demixing line, mixtures form a system of slowly moving quasi-particles with the interaction rapidly decreasing at large distances. In this limit, the scattering of quasi-particles reduces effectively to s-wave scattering.^{22,23} Then all the interaction processes can be described using only one parameter, namely the s-wave scattering length a_0 . The value of a_0 has to be obtained experimentally. For ³He-⁴He mixtures, the interaction between ³He quasi-particles is effectively attractive, so the s-wave scattering length is negative, $a_0 = -0.088nm$.²⁴ The interaction term of Eq. (4) is given by²⁵

$$\varepsilon_{\rm int} = \frac{p_F^2}{2m_i} \left[\frac{4}{3} \lambda + \frac{4}{15} \left(11 - 2 \ln(2) \right) \lambda^2 + \cdots \right]$$
(6)

where the dimensionless parameter $\lambda = p_F a_0 / \pi \hbar \propto n_{34}^{1/3}$ is assumed to be small, which is equivalent to the condition $n_{34} \ll |a_0|^{-3}$, and the single-particle *impurity* effective mass m_i is defined as

$$m_i = \lim_{n_{34} \to 0} m^*(n_{34}) \tag{7}$$

The Solubility of Spin-Polarized ³He in ⁴He

while the effective mass caused by quasi particle interactions is

$$\frac{m^*}{m_i} = 1 + \frac{8}{15} \lambda^2 (7 \ln 2 - 1) + \cdots$$
(8)

3. The local effective potential model of Bardeen, Baym and Pines,¹¹ where it is assumed that the quasi-particle interaction can be described by an effective potential $v_{\text{eff}}(\vec{q}) = v_0 \cos(q/k_s)$, where k_s is the effective range of the potential with strength v_0 . The lowest Legendre moments of the scattering amplitudes were calculated by van de Haar, Frossati and Bedell¹² (hereafter referred to as HFB) in the thermodynamic limit.⁽¹⁾ An attempt to obtain better agreement with experiment was made by defining $k_s(n_{34}) = k_s(1 + \beta n_{34}/n_s)$, with β a dimensionless parameter allowing for a stronger dependence of v_{eff} on n_{34} . With the interaction parameters we can obtain the Fermi liquid interaction parameter $f_{\vec{p}\vec{p}'}^{\sigma\sigma'} = f_{\vec{p}\vec{p}'}^s + \sigma\sigma' f_{\vec{p}\vec{p}'}^a$, with $|p| = |p'| = p_F$. The moments of $f_{\vec{p}\vec{p}'}^s$ enter the expression for ε_{int} at T = 0 as follows,²⁶

$$\varepsilon_{\text{int}} = 2 \int_{0}^{p_{F}} (f_{\vec{p}\vec{p}'}^{s})_{|p| = p_{F}} \frac{d^{3}p'}{(2\pi\hbar)^{3}}$$

$$= v_{0} n_{34} \left(1 + \frac{9}{4} \frac{1 - \cos(2K)}{K^{4}} + \frac{3}{2} \frac{\cos(2K)}{K^{3}} + 3 \frac{\sin(2K)}{K^{2}} + \cdots \right)$$
(9)

where $K = k_F / k_s = p_F / \hbar k_s \propto n_{34}^{1/3}$. In this approach, m^* is defined as

$$\frac{m^*}{m_i} = 1 + \frac{F_1^s}{3} \tag{10}$$

where the Landau parameter F_1^s can be calculated from the scattering amplitude a_1^s given in Ref. 12.

In principle, similar calculations can be done using other model representations for the effective potential (see, e.g., reviews^{25,27}).

¹Note a misprint just after their equation 24 d):

$$\beta_0 = 9\alpha_0 \left(\frac{1}{K^2} - 2\right) + \nu_0 \frac{1}{K} \left(\frac{9}{K} - 12S_K\right)$$
$$S_K = \sin(2K) \quad \text{and} \quad C_K = \cos(2K)$$

| | TA | BL | Æ | I | |
|--|----|----|---|---|--|
|--|----|----|---|---|--|

Mixture Related Experimental Constants. The Digits Given in Brackets are Given for Smoothness Only.

| | P = 0 MPa | P = 1 MPa | P = 2 MPa |
|------------------------------------|--------------------------|--------------------------|--------------------------|
| m_i/m_3 | 2.34 | 2.64 | 2.89 |
| $n_3 (\text{mol/m}^3)$ | 271(20) | 329(89) | 361(15) |
| v_{34}^{m} (m ³ /mol) | $28.0(97) \cdot 10^{-6}$ | $25.6(67) \cdot 10^{-6}$ | $24.0(96) \cdot 10^{-6}$ |
| $x_{s}(\%)$ | 6.65 | 9.49 | 8.55 |
| μ_3 (J/mol) | -20.56(15) | +12.76(13) | +42.02(61) |
| $\varepsilon_0 (J/mol)$ | -23.09(64) | +9.82(29) | +39.17(48) |
| | | | |

2.2. Comparison of the Expansions in Density

In Table I, we have given all the experimental mixture-related parameters used in this work. In both the s-wave approach and the theory of HFB, there are parameters which are not known to very high accuracy. In the s-wave picture, we took (for zero pressure) the most up-to-date value for a_0 , -0.088nm, and a more than twice larger value, -0.180nm, which gives a more accurate number for the experimental value of the demixing concentration,²¹ see Table II. However, one should always keep in mind that the s-wave scattering approximation rapidly loses its accuracy when the concentration exceeds several percent.^{25,27} At concentrations close to x_s , the s-wave approximation works qualitatively in thermodynamics, but becomes very unreliable for numerical estimates of T_c . There are indications²⁸ that the effective interaction at higher concentrations is weaker than the extrapolation of the low-concentration results in the s-wave approximation.

On the other hand, the HFB theory is formally applicable at all concentrations, but operates with a model interaction potential with several fitting parameters. It is unclear how reliable the use of the same model potential for helium mixtures at different polarizations and with different densities of states is.

| Scattering Lengths Used in s-wave Calculations | | | | |
|--|--------------------|---------------|--------------------|--|
| | P = 0 MPa | P = 1 MPa | P = 2 MPa | |
| Set 1: a_0 (nm) Set 2: a_0 (nm) | $-0.180 \\ -0.088$ | -0.166 -0.088 | $-0.138 \\ -0.088$ | |

TABLE II

| TILD Model Fit Farameters | | | | |
|---------------------------|------------------------|------------------------|-------------------------|--|
| | P = 0 MPa | P = 1 MPa | P = 2 MPa | |
| $v_0 ({\rm Jm^3})$ | $-1.39 \cdot 10^{-51}$ | $-1.10 \cdot 10^{-51}$ | -0.96 · 10 ⁻ | |
| $c_s (m^{-1})$ | $4.0 \cdot 10^{9}$ | $4.4 \cdot 10^{9}$ | $5.8 \cdot 10^{9}$ | |
| 3 | 0.0 | 0.5 | 0.7 | |

TABLE III

 $\frac{v_0 \text{ (Jm}^3)}{k_s \text{ (m}^{-1})} \xrightarrow{-1.39 \cdot 10^{-51}}{4.0 \cdot 10^9} \xrightarrow{-1.10 \cdot 10^{-51}}{4.4 \cdot 10^9} \xrightarrow{-0.96 \cdot 10^{-51}}{5.8 \cdot 10^9}$ Instead of comparing many different interaction potentials, we decided to show results on one particular set of parameters for one particular ntial, namely the one given by HFB with the parameters given in

HFB Model Fit Parameters

Instead of comparing many different interaction potentials, we decided only to show results on one particular set of parameters for one particular potential, namely the one given by HFB, with the parameters given in Table III. This set of parameters is essentially the fit made by HFB to magnetostriction data,¹² with the range parameter k_s adjusted to reproduce the experimental value of the solubility limit. Using the original fits of HFB did not change our results for x_s by more than 5%.

It is instructive to compare the different expansions of the chemical potential as a function of the density. The experimental data and the HFB-expansion show the same functional dependence of μ_{34} on n_{34} , Eq. (5). The s-wave picture gives rise to an additional term proportional to $n_{34}^{4/3}$. In Table IV, we compare the different terms for P = 0 MPa. The model of HFB is in good agreement with the experimental fit. The lowest order terms of the s-wave approach, especially for $a_0 = -0.088$ nm, also agree with the experimental data. Because the functional dependence on n_{34} is different, it is better to compare the chemical potentials graphically, see Fig. 1, where we plotted chemical potential curves at P = 1 MPa for several models.

The steepest curve is the one for a Fermi gas without any interaction, which is equivalent to taking the $a_0 \rightarrow 0$ limit in the s-wave model or taking the $v_0 \rightarrow 0$ limit in the HFB model. At concentrations up to a few percent, the Fermi gas model and the s-wave model with $a_0 = -0.088$ nm fit the

TABLE IV

Comparison Between Different Expansions of the Chemical Potential of ³He in Mixture μ_{34} (J/mol) at P = 0 MPa in the ³He Density n_{34} (10⁶ mol/m³)

| | $\propto n^{2/3}$ | $\propto n$ | $\propto n^{4/3}$ | $\propto n^{5/3}$ |
|-----------------------------------|-------------------|-------------|-------------------|----------------------|
| experimental | 194.6 | -201 | 0 | $-1.04 \cdot 10^{4}$ |
| s-wave, $a_0 = -0.18 \text{ nm}$ | 194.4 | -387 | 1110 | not listed |
| s-wave, $a_0 = -0.088 \text{ nm}$ | 194.4 | -189 | 266 | not listed |
| HFB model | 181.9 | -234 | 0 | $-1.10 \cdot 10^{4}$ |



Fig. 1. The difference in ³He chemical potential between pure ³He and ³He in ³He.⁴He mixtures as function of concentration at P = 1 MPa. Full line: experimental curve, dotted line: s-wave approximation with $a_0 = -0.088$ nm, dash-doubly dotted line: free Fermi gas, dotted line: s-wave approximation with $a_0 = -0.166$ nm, dash-dotted line: HFB model.

experimental data best, while the other curves lie below experiment. At higher concentrations, up until 15%, deviations are smallest for the HFB model and the s-wave model with $a_0 = -0.166$ nm, while the ones mentioned before keep on increasing steeply. This behaviour is seen at all pressures.

At concentrations higher than x_s , the HFB curve is less steep than the s-wave curves. This means that in the s-wave picture it is more difficult to increase the concentration, because the energy needed to do so is higher.

The experimental data are extrapolated from 25–700 mK to T=0 (see Ref. 21, and references therein). This means it is also possible to plot the chemical potential of mixtures of concentration above the T=0 solubility limit $(x_s(T) = x_s(0)(1 + \beta_T T^2))$, up to X = 16%. As we will see in Sec. IIIE, the solubility limit can be as high as 25%; theoretical curves lose their accuracy when extrapolated to such high concentrations.

Recent analysis²⁸ of experimental data^{29,30} on spin dynamics in ³He-⁴He mixtures provides an additional confirmation that the effective attraction between ³He particles at high concentration is strongly suppressed in comparison with data at low concentration.

3. SOLUBILITY IN FINITE MAGNETIC FIELD

3.1. The Relative Amount of ³He

In a magnetic field *B*, the total system, the concentrated phase and the dilute mixture will have polarizations Δ , Δ_3 and Δ_{34} , respectively, where

$$\Delta = \frac{N_3 \Delta_3 + N_{34} \Delta_{34}}{N_3 + N_{34}} \tag{11a}$$

$$\Delta_i = \frac{N_i^{\uparrow} - N_i^{\downarrow}}{N_i^{\uparrow} + N_i^{\downarrow}} \tag{11b}$$

Here \uparrow (\downarrow) stands for spin-up (down) particles and the index *i* is either "3" or "34". The relation between polarization and magnetic field is given by the susceptibility

$$\chi = \mu_0 \frac{\partial m}{\partial B} = \mu_0 \mu_{3N} n \frac{\partial \Delta}{\partial B}$$
(12)

where $\mu_0 = 4\pi \cdot 10^{-7}$ J/(Am²) is the magnetic permeability of the vacuum, $\mu_{3N} = 0.778$ mK/T the ³He nuclear magnetic moment and *n* is the number density of ³He particles. The susceptibility has only been measured in thermal equilibrium at low field; in the pure phase the measurements have been done by Ramm *et al.*^{31,32} and in mixtures by Ahonen *et al.*³³ Using the rapid melting technique, the susceptibility of pure ³He along the (depressed) melting curve has been measured by Wiegers, Wolff and Puech,³⁴ for polarizations up to $\Delta = 0.6$. Otherwise, we have to estimate the susceptibility theoretically.

There is a distinct difference between polarization obtained in thermodynamic equilibrium by applying a large external field (the so-called *brute force technique*) and polarization determined by the history of the system, as in rapid melting experiments. This non-equilibrium polarization can be very long-lived,^{7,8,35,36} and the evolution of the system, in this case, takes place at constant *total* polarization Δ , Eq. (11a). This imposes an additional constraint on the thermodynamic variables Δ_3 and Δ_{34} . In such a situation, one should describe the distribution of the ³He particles *and* polarization between the pure and dilute phase as a function of overall polarization Δ , the total number of ³He particles in the system, $N = N_3 + N_{34}$ and the total number of ⁴He particles N_4 . As a result, the solubility limit x_s at T = 0 and the polarizations Δ_3 and Δ_{34} of both phases depend on Δ , N and N_4 . Then the corresponding value of $x_s(\Delta, N, N_4)$ differs from the thermodynamic equilibrium value $x_s(\Delta_{34}, N_{34}/N_4)$ which does not depend on N/N_4 . Therefore, one should carefully specify the experimental conditions under which the system demixes into a pure and a dilute phase.

In a magnetic field, the Gibbs-Duhem relations, Eqs. (3) change to

$$dP_3 = n_3 \, d\mu_3 + m_3 \, dB \tag{13a}$$

$$dP_{34} = n_{34} \, d\mu_{34} + n_4 \, d\mu_4 + m_{34} \, dB \tag{13b}$$

where $m = \mu_{3N} n\Delta$ is the magnetization and *B* the magnetic field. Although Eqs. (13) are coupled by the condition $dP_3 = dP_{34}$ in order to keep the pressure equal in both subsystems, many parameters can change in reaction to a change in, say, the external magnetic field.

A possible simplification is to restrict oneself to certain limits of X. For example, in the limit of $X \rightarrow 1$, when most of the system is in the pure phase and the amount of mixture is very small, the presence of the dilute phase with polarization Δ_{34} and concentration x_s practically does not affect the volume of the pure phase, nor its polarization $\Delta_3 \approx \Delta$, which is constant. Then the polarization Δ_{34} and the ³He density n_{34} depend only on Δ_3 . The calculation of $x_s(\Delta_3)$ for these conditions is performed in Sec. IIIB. The opposite case, leading to a different solubility, corresponds to a system with a small amount of pure phase. In this case $\Delta_{34} \approx \Delta$ is constant, and the polarization Δ_3 is a function of Δ_{34} while x_s depends only on Δ_{34} and is not sensitive to Δ_3 .

Another possible approach is to reduce the number of parameters by demanding that certain boundary conditions must be met. In Sec. IIIC for example, we restrict ourselves to the case where $d\mu_4 = 0$ and $dP_3 = dP_{34} = 0$. Both approaches limit themselves to specific experimental arrangements, which makes it *a priori* difficult to compare them to each other.

3.2. $X \uparrow 1$

The polarization dependence of the chemical potential μ_{34} was calculated by Bashkin and Meyerovich²⁵ in the s-approximation using second order perturbation theory. In general, if the *total* polarization of the system is constant, the chemical potentials of the ³He Fermi liquids in the pure and dilute phases at T=0 can be written as

$$\mu_{3}^{\sigma}(n_{3},\Delta) = \frac{(3\pi^{2}n_{3})^{2/3}}{2m_{3\sigma}^{*}(\Delta,n_{3})} (1+\sigma\Delta)^{2/3}$$
(14a)

$$\mu_{34}^{\sigma}(n_{34}, \Delta) = -\varepsilon_0(\Delta, \mu_4, n_{34}) + \frac{(3\pi^2 n_{34})^{2/3}}{2m_{34,\sigma}^*(\Delta, n_4, n_{34})} (1 + \sigma \Delta)^{2/3} \quad (14b)$$

If we assume, as was suggested in the previous Section, that the amount of ⁴He in the system is very small and the system consists mostly of pure ³He, $X \approx 1$, $\Delta_3 \approx \Delta$, we can equate $\mu_3^{\sigma}(\Delta)$ and $\mu_{34}^{\sigma}(\Delta)$ without having to use further equations for the polarizations. Restricting ourselves to the simplest realistic approximation in which ε_0 is a constant and the effective masses do not depend on the signs of their spins, we get

$$x_{s}(\Delta) = \sum_{\sigma=\uparrow,\downarrow} \left[\left(\frac{x_{s}(0)}{2} \right)^{2/3} + \xi_{\sigma}^{2/3} \right]^{2/3}$$
(15a)

$$\xi_{\sigma} = \left(\frac{1}{2} n_3 v_{at}\right)^{2/3} \left[(1 + \sigma \Delta)^{2/3} \frac{m_{34}^*(\Delta)}{m_3^*(\Delta)} - \frac{m_{34}^*(0)}{m_3^*(0)} \right]$$
(15b)

where $x = v_{al}n_{34}$, with $v_{al}n_{34}$, with $v_{al} = 1/(n_{34} + n_4)$ the average volume per atom in the mixture. This approximation is quite accurate for mixtures of not very high concentration; the error caused by the assumption $m_{\uparrow}^* \simeq m_{\downarrow}^*$ is unknown and should be higher than the error caused by the assumption $\varepsilon_0 = \text{constant.}$

3.3. $X \downarrow x_s$

Imposing the conditions $d\mu_4 = 0$ and dP = 0 on Eqs. (13) gives us the simple relation¹

$$d\mu_i = -\mu_{3N} \Delta_i \, dB \tag{16}$$

where $i \in \{3, 34\}$. Integrating this at constant density n_i gives

$$\mu_i(n_i, B) - \mu_i(n_i, 0) = -\mu_{3N} \int_0^B \Delta_i(B') \, dB' \equiv -I_i(n_i, B) \tag{17}$$

Then we set $n_{34} = n_s(B)$, the saturation density of the mixture in a magnetic field *B* and we use the phase equilibrium condition $\mu_3(n_3, B) = \mu_{34}(n_s(B), B)$ for both B = 0 and finite *B* which, combined with Eq. (17) and i = 3, 34 gives

$$\mu_{34}(n_s(B), 0) - \mu_{34}(n_s(0), 0) = I_{34}(n_s(B), B) - I_3(n_3, B)$$
(18)

¹ Allowing μ_4 to vary would lead to a more general result than the one presented here, where we keep it fixed. One could e.g. write

$$d\mu_4 = \left(\frac{\partial\mu_4}{\partial n_{34}}\right)_p \left(\frac{\partial n_{34}}{\partial \mu_{34}}\right)_p d\mu_{34} = F(n_{34}) d\mu_{34}$$

Then Eq. (16) would read

$$(1 + F(n_{34})) d\mu_{34} = -\mu_{3N} \Delta_{34} dB$$

with $F(n_{34}) < 0$. Thus including the osmotic pressure effect tends to decrease the solubility enhancement. We thank Dr. G. Vermeulen for pointing this out to us.

With the help of the magnetic susceptibility we change from magnetic field to polarization, and use the identity

$$I(n, B) = \mu_0 \mu_{3N}^2 n \int_0^{\Delta(B)} \frac{\Delta'}{\chi(\Delta')} d\Delta'$$
(19)

in order to calculate the limiting density $n_s(B)$ self consistently from Eq. (18). In this approach, we need to know only the zero field chemical potential of the dilute phase, and do not need to know the chemical potential of the pure sub-system at all. In this calculation, the ratio of ³He to ⁴He particles enters the result only if we try to calculate the concentration x instead of the density n_{34} . Here we assume that we have the same relation between x and n as in the unpolarized case, which is only true in the case $X \rightarrow x_s$.

3.4. Low Polarization (Thermodynamic Limit)

Phase separated mixtures have a rather low susceptibility, $^{31-33}$ which results in a very low polarization when placed in experimentally available fields (brute force polarization). For this reason, we derive approximations for the limit of low polarization, where the system is in thermodynamic equilibrium with an external magnetic field. Expanding the Gibbs function in the magnetization yields^{37,38}

$$G(\Delta) = G(0) + \frac{1}{3}NE_F G_m \Delta^2 - \mu_{3N} NB\Delta$$
⁽²⁰⁾

The second term on the right side of the last equation represents the change of pressure of the system due to the magnetization (dPV/dM) at constant density, while the third term is the Zeeman energy. The magnetic Grüneisen parameter G_m is defined as

$$G_{m} = \left(\frac{\partial \ln E_{F}^{**}}{\partial \ln n_{3}}\right)_{\mu_{4}}$$
$$= 1 - \left(\frac{n_{3}}{\chi}\frac{\partial \chi}{\partial n_{3}}\right)_{\mu_{4}} = 1 + (1 - x)\left(\frac{x}{\chi}\frac{\partial \chi}{\partial x}\right)_{\mu_{4}}$$
(21)

with $E_F^{**} = (1 + F_0^a) E_F^* = ((1 + F_0^a)/(m^*/m)) E_F$. G_m was calculated for helium mixtures by Bedell and van de Haar.³⁹

From Eq. (20) it follows³⁷ that the limiting solubility $x_s(B)$ can be expanded as

$$x_s(B) = x_s(0)(1 + \beta_M B^2 + \cdots)$$
 (22)

with

$$\beta_{M} = \frac{1}{2\mu_{0}} \frac{(1-x)^{2}}{xn_{4}} \left(\frac{\partial \mu_{34}}{\partial n_{34}}\right)^{-1} \left(\frac{\chi_{34}}{n_{34}} A - \frac{\chi_{3}}{n_{3}}\right)$$
(23)

In the calculation of Dalfovo and Stringari,³⁷ which includes the magnetostriction effect, A is equal to the magnetic Grüneisen parameter, G_m . To compare with this thermodynamic calculation, we expand the integral given in Eq. (19) in the magnetic field:

$$I(n, B) = \frac{1}{2\mu_0} \frac{\chi}{n} B^2 + \dots$$
 (24)

which, although the present work does not include magnetostriction, leads to an equation for β_M of the same form as Eq. (23), with A = 1.

Expanding Eq. (15) in the magnetic field and making the further assumption $m_i^*(\Delta) = m_i^*(0)$ leads to

$$\beta_M = \frac{\alpha}{3x} \left[\alpha \left(\frac{x}{2} \right)^{-1/3} - \left(\frac{x}{2} \right)^{1/3} \right] \left(\frac{\Delta_3}{B} \right)^2$$
(25a)

$$\alpha = \frac{m_{34}^*}{m_3^*} \left(\frac{n_3}{n_{34} + n_4}\right)^{2/3} \tag{25b}$$

In the limit of a gas of Fermi particles with mass m^* and susceptibility

$$\chi_{FL} = \mu_0 \mu_{3N}^2 \frac{3}{2} \frac{n}{E_F} \frac{m^*/m}{1 + F_0^a}$$
(26)

all models give similar expressions for β_M . This is to be expected, since they all agree on the first terms of the expansion Eq. (4) of μ_{34} in n_{34} ($\varepsilon_{int} = 0$):

$$\beta_{M} = \frac{3}{8} \mu_{3N}^{2} \frac{1}{E_{F,3}^{*} E_{F,34}^{*}} \left[\frac{m_{34}^{*}}{m_{3}^{*}} \left(\frac{n_{3}}{n_{34}} \right)^{2/3} A_{34} - A_{3} \right]$$
(27)

with $E_{F,3(34)}^*$ the Fermi temperature of the pure (mixture) system using the effective mass m_i^* .

The values of A_3 and A_{34} are summarized in table V. It should be kept in mind though, that according to Fig. 1, the Fermi gas limit is a poor approximation for the chemical potential. Solving Eq. (18) with the approximation (24) gives A = 1, while the thermodynamic expansion (taking the concentration derivative of a not well known mixture susceptibility) yields $A = G_m = 0.53$. For a Fermi gas, including the magnetostriction effect gives $A = G_m = \frac{2}{3}$.

For the result obtained using Eqns. (18–19), the sign of β_M is determined by the sign of $(\chi_{34}/n_{34} - \chi_3/n_3)$, which changes sign between 0 and 1 MPa. At P = 0, A > 0.78 ensures $\beta_M > 0$, explaining the difference in sign between the thermodynamic expansion and the model for $X \downarrow x_s$. At high

L. P. Roobol et al.

TABLE V

The Value of A_3 and A_{34} , Defined in Eq. (27). All Coefficients are Obtained in the Limit of Low Polarization Using a Fermi Liquid Susceptibility

| | X↑1 Eq. (25a) | thermodynamic expansion | $\begin{array}{c} X \downarrow x_s \\ \text{Eq. (24)} \end{array}$ |
|-----------------------------------|---|---|--|
| K ₃ K ₃₄ | $\frac{1/(F_0^a(3))^2}{1/(F_0^a(3))^2}$ | $\frac{3(1-x)/(F_0^a(3))}{3G_m(1-x)/(F_0^a(34))}$ | $\frac{3(1-x)/(F_0^a(3))}{3(1-x)/(F_0^a(34))}$ |

pressure the parameters change such that both approaches agree on the sign again.

In the case of $X \uparrow 1$, the sign of β_M depends only on the ratio of the effective masses and the densities, which at all pressures yields a positive sign.

In table VI, we have listed the values of β_M at three different pressures using the different chemical potentials calculated with the various models, with the susceptibilities taken from experiment.³¹⁻³³

3.5. Results for Arbitrary Polarization

In Fig. 2, we plot the limiting solubility x_s as a function of total polarization as calculated from Eq. (15). Because we here consider the case $X \uparrow 1$, the variable of interest is $\Delta_3 \approx \Delta$. As it turns out, there are only minor differences between the two sets of scattering lengths given in Table II. In this calculation we have assumed that the effective masses as well as ε are independent of polarization and equal to their zero field values, effectively reducing it to a free Fermi gas approximation. Depending on the pressure, the solubility limit is enhanced by a factor 2–4. Also, the zero pressure curve intersects the other two, indicating that one should be in the right pressure range if one aims for the highest solubility at a given polarization. Though at high concentrations ($x_s(\Delta = 1) \approx 0.2$) the s-wave model loses its

| T. | A | BL | Æ | VI |
|----|---|----|---|----|
| _ | | _ | | |

The Coefficient of the Quadratic Term in the Expansion (22), β_M

| | 0 MPa | 1 MPa | 2 MPa |
|--|--|---|---|
| $X \uparrow 1, \text{Eq.} (25a)$ $X \downarrow x_s, \text{Eq.} (24)$ Dalfovo and Stringari | $9.9 \cdot 10^{-7} \\ 1.9 \cdot 10^{-6} \\ -1.5 \cdot 10^{-6}$ | $1.1 \cdot 10^{-6} \\ -4.8 \cdot 10^{-6} \\ -4 \cdot 10^{-6}$ | $5.4 \cdot 10^{-7} \\ -3.5 \cdot 10^{-5} \\ -6 \cdot 10^{-6}$ |



Fig. 2. The limiting solubility x_s as function of polarization $\Delta \approx \Delta_3$, for a Fermi gas, in the limit $X \approx 1$. Full line: P = 0 MPa, dashed line: P = 1 MPa, dotted line: P = 2 MPa.

accuracy and cannot be applied quantitatively, it, nevertheless, indicates that the solubility may experience dramatic changes with polarization. The densities and zero field solubility were taken from Ref. 21, while the effective masses were taken from Ref. 40 (pure phase) and Ref. 33 (mixture).

Fig. 3, which plots the limiting solubility as function of the polarization of the dilute phase, is the result of solving Eq. (18) numerically for the density $n_{34}(\Delta)$. In this case $(X \downarrow x_s)$ the mixture polarization $\Delta_{34} \approx \Delta$ is the proper variable to use. For χ_3 , the susceptibility of the pure phase, we used the nearly metamagnetic model (an extension to finite polarization of Fermi liquid theory), put forward by Bedell and Sanchez-Castro¹³ and worked out in the case of liquid ³He by Sanchez-Castro, Bedell and Wiegers.¹⁴ For the properties of the dilute phase, we used the potential model of van de Haar, Frossati and Bedell,¹² as discussed above. Because the susceptibility is a function of ³He density, is necessary to evaluate the integral I_{34} , Eq. 17 each time the density is changed, which makes the computation rather time consuming. In order to calculate the concentration xfrom the density n_{34} , we assume that the relation between density and concentration is independent of polarization. We also see that in this model



Fig. 3. The limiting solubility x_s as function of polarization $\Delta \approx \Delta_{34}$, as calculated using the HFB model for the mixture and the nearly metamagnetic model for the pure phase, in the limit $X \approx x_s$. Full line: P = 0 MPa, dashed line: P = 1 MPa, dotted line: P = 2 MPa.

the zero pressure curve intersects the P = 1 MPa curve, but that the changes in solubility limit are not nearly as dramatic as in the s-wave model. Only the P = 0 result shows an increase of 30% in x_s at $\Delta = 0.95$, whereas the curves at 1 and 2 MPa go through a shallow minimum, approaching more or less the zero field value at full polarization.

In Fig. 4 we plot the polarizations Δ_3 and Δ_{34} as a function of magnetic field, *along the demixing line* at a pressure of 1 MPa. Because the susceptibility of the nearly metamagnetic model goes through a maximum at $B \approx 10^2$ T,¹⁴ the slope of the polarization curve decreases, while the polarization of the dilute phase continues to increase more or less linearly in the HFB model. This results in the fact that at $\Delta \propto 0.7$, the polarization curves intersect and at higher magnetic fields, $\Delta_{34} > \Delta_3$.

Fig. 5 shows the magnetic energy stored in the mixture and the pure phase as function of magnetic field B at P = 1 MPa. The energy shift is almost 2 J/mol at $\Delta = 1$, which is to be compared with Fig. 1. The difference between the curves (see Eq. (18)) however is quite small, making the sign of $(x_s(\Delta) - x_s(0))$ not very certain, but we can conclude that at



Fig. 4. The polarizations Δ_3 and Δ_{34} in the dense and dilute phase, respectively, as a function of magnetic field *B* at a pressure of 1 MPa along the demixing line, using the HFB model for the mixture and the nearly metamagnetic model for the pure phase. Full curve: Δ_3 , dashed curve: Δ_{34} .

elevated pressures this model predicts the change in limiting density to be small.

4. THE SUPERFLUID TRANSITION TEMPERATURE

Superfluid ³He dissolved in superfluid ⁴He is one of the "Holy Grails" of (ultra-) low temperature physics. It would provide us with a unique binary mixture of different superfluids which has never been observed before. As was mentioned in the introduction, different theoretical models give vastly different predictions not only for the transition temperature itself, but also for the type of superfluid pairing. Whatever the mechanism, the superfluid transition temperature can be enhanced by either enlarging the interatomic (attractive) interaction, or by increasing the density. We now, for the first time, calculate T_C taking into account that the ³He density in the mixture is affected by the polarization.



Fig. 5. The magnetic energies I_3 and I_{34} stored in the pure and the dilute subsystem, respectively, at a pressure of 1 MPa, along the demixing line. Full line: pure phase, dashed line: mixture.

4.1. S-wave Pairing in the Weak Coupling Limit

The interaction between ³He particles in the s-wave channel is attractive, making the s-wave superfluid transition one of the most probable scenarios for dilute mixtures.^{12,27} In the case of s-wave pairing, the pairing particles have opposite spin projections. The spin polarization makes the Fermi momenta of particles with opposite spins, p_{\uparrow} and p_{\downarrow} , different from each other. If the polarization is low, s-wave pairing is still feasible, though the corresponding Cooper pairs have non-zero momentum, while the transition temperature $T_C(\Delta_{34})$ is much lower than the transition temperature $T_C(0)$ in the absence of polarization. When the polarization becomes higher so that the difference of the Fermi energies for up- and down spins $p_{\uparrow}^2/2m_{\uparrow}^* - p_{\downarrow}^2/2m_{\downarrow}^*$ becomes larger than the pairing energy ($\propto T_C$), s-wave pairing becomes impossible. In other words, to expect standard s-wave superfluidity, the polarization should not exceed $T_C(\Delta=0)/T_F$. Since according to experimental data the value of $T_C(0)$ is much less than 1 mK, even a polarization Δ_{34} less than 1% makes s-wave pairing impossible.

4.2. Standard *p*-wave Pairing in the Weak Coupling Limit

If the pairing particles have equal spin projections, the splitting into two Fermi spheres is not crucial since the onset of superfluidity first occurs on the larger sphere. The radius of this Fermi sphere increases because of polarization and could also increase if the ³He density is enhanced. This phase will have properties similar to the superfluid A_1 phase of pure ³He in high magnetic field.²⁷ The expression of T_C can be written as

$$T_C^{\sigma} \simeq T_F^{\sigma} \exp(-\gamma/x) \tag{28}$$

where $\gamma \approx b/a_0$ characterizes the strength of the interaction in the p-wave channel, $T_F^{\sigma} = T_F (1 + \sigma \Delta)^{2/3}$. Since there is insufficient information on the interaction in the p-wave channel, it is very difficult to give a reliable estimate for γ . This makes the numerical estimates for T_C (Eq. 28) not very accurate even by order of magnitude especially because γ enters the index of the exponent and $\gamma/x \ge 1$.

In the approach of Ref. 41,

$$\gamma = 3x/N^{\sigma}(0) a_1^{\sigma\sigma} \tag{29}$$

where $N^{\sigma}(0)$ is the density of states for particles with spin σ , and $a_1^{\sigma\sigma}$ is one of the Fermi liquid scattering parameters. The maximal possible density of states, and therefore, the highest transition temperature is determined by the maximal solubility of ³He in ⁴He.

The results for T_C with γ (Eq. 29) calculated with the help of the HFB model are given in Fig. 6. We see an enormous increase in T_C , from 2 to 4 orders of magnitude, depending on the pressure, with polarization. Unfortunately, the highest increase in T_C is predicted for P = 0, which gives a far lower T_C than in the case P = 1 MPa. The estimate for 1 MPa, starting at a modest 10 μ K, approaches the millikelvin regime at full polarization.

In Fig. 7, we plot the ratio of the T_c computed in this paper to the T_c 's found in the original paper by HFB,¹² when the change in solubility was not taken into account. We see that including the polarization dependence of the solubility enhances the value of T_c by a factor of 4 only. In the original paper of HFB, keeping the solubility constant, the polarization increases the transition temperature by orders of magnitude. This is the effect of the change in effective interaction and the enlargement of the up-spin Fermi sphere, which in this model appears to be the main reason for the large enhancement of T_c .



Fig. 6. The p-wave superfluid transition temperature along the demixing line as function of polarization $\Delta \approx \Delta_{34}$ at different pressures, as calculated with the HFB/nearly metamagnetic model. The change in ³He density with polarization is taken into account. Full line: P = 0 MPa, dashed line: P = 1 MPa, dotted line: P = 2 MPa.

In the dilute gas model, the p-wave scattering amplitude b and, therefore, γ should be considered as constants independent of polarization. Then the effect of spin polarization on T_c reduces solely to the concentration dependence in Eq. (28). The value of b should be determined from experimental data. Unfortunately, the existing experiments at low concentration cannot provide the value of b; experimental information on b can only be obtained from experiments at high polarizations.²⁷ Therefore, at present the value of b is completely unknown, and it is impossible to give a numerical estimate for T_c .²⁸

We can use Eq. (29) to fix γ in Eq. (28) at zero polarization. Doing so, we can use this last equation to compute the p-wave superfluid transition temperature with the limiting concentrations resulting from Eq. (15). We have plotted the values of T_C obtained in this way in Fig. 8. We see a dramatic increase in T_C of up to 5 orders of magnitude, generated by the large increase in $x_s(\Delta)$. As in the case where $X \downarrow x_s$, the largest increase is seen at P = 0, while the highest T_C is predicted for P = 1 MPa. Note that in this approach the interactions are kept constant, while the increase in T_C



Fig. 7. The enhancement of the superfluid transition temperature along the demixing line compared to the HFB model (which keeps x_s constant) at different pressures. Full line: P = 0 MPa, dashed line: P = 1 MPa, dotted line: P = 2 MPa.

is generated solely by the increase in density. Making a plot like Fig. 7 would thus yield the trivial result $T_c(\Delta, x_s(\Delta)) = T_c(\Delta, x_s(0))$.

4.3. Kohn-anomaly p-wave Pairing in the Weak Coupling Limit

Recently, Kagan and Chubukov⁴² have shown that the effective interaction in the p-wave channel is dominated by the attraction resulting from renormalizations induced in the second order terms of the s-wave interaction. In the weak coupling limit this approach yields higher values of the transition temperature than the standard p-wave model discussed above, but the dependence on concentration is weaker:

$$T_C^{\sigma} = T_F^{\sigma} \exp(-\bar{\gamma}/x^{2/3}) \tag{30}$$

where the dimensionless constant $\bar{\gamma} \approx 1$. Since at low concentrations $x^{2/3} \ll x$, the increase in T_C as a result of the enhanced solubility x is not as spectacular as in Eq. (28). However, the absolute value of T_C^{30} itself should be much larger than the values obtained in the standard p-wave pairing model.²⁸



Fig. 8. The p-wave superfluid transition temperature using the polarization dependent limiting solubility calculated for a Fermi gas with $X \approx 1$, at different pressures. $T_{c}(0)$ was taken from Eq. (29) and is therfore the same as in Fig. 6. Full line: P = 0 MPa, dashed line: P = 1 MPa, dotted line: P = 2 MPa.

5. CONCLUSIONS

We have applied some theoretical approaches to spin-polarized ³He and ³He-⁴He mixtures to the problem of the solubility limit. For low polarization, all calculations agree that changes in the solubility will be small, although there is disagreement even on the sign of the effect.

We examined two approaches which apply at arbitrary polarization, the HFB model combined with the nearly metamagnetic model in the limit $X \downarrow x_s$, and the dilute gas model in the limit $X \uparrow 1$. At higher polarizations, these models disagree: for $X \approx x_s$ the solubility is predicted to go through a minimum at P = 1 - 2 MPa, after which it increases to about the zero polarization value. At zero pressure, n_{34} is predicted to increase by 30%. This is certainly not as optimistic as the model applied to the case $X \approx 1$, which predicts a 2-4 fold increase at all pressures (however, as it was mentioned before, the s-wave approach loses its accuracy with an increase in concentration).

Because the ³He superfluid phase transition temperature is a monotonously increasing function of n_{34} , the solubility limit sets an upper

360

limit for T_C . Both theories predict an enormous enhancement of T_C with polarization, though the sources of this increase are different. In the s-wave approximation (or, better, dilute gas model) the effective interaction is considered constant while all observable changes caused by polarization are explained in terms of relative changes in densities of state near the Fermi surfaces for up- and down spins. Then the increase in transition temperature is explained by a considerable increase in the limiting solubility and, therefore, increase in the density of states without changing the effective interaction. The HFB model also gives a large enhancement in T_C , about the same order of magnitude. However, this model relates all the effects to changes in phase space and effective interaction with polarization, while the changes in density play a secondary role. If there is a density (solubility) effect in the HFB/nearly metamagnetic model, it tends to diminish T_C , except at zero pressure.

Note, that the results can be very sensitive to whether one deals with equilibrium polarization caused by an external magnetic field, or with a long-lived non-equilibrium polarized state with a given overall polarization. In the latter case, which corresponds to more realistic ways of obtaining high polarizations, the calculations of the limiting solubility and polarization should be done with the additional constraint that the *total* polarization $\Delta = (\Delta_3 N_3 + \Delta_{34} N_{34})/(N_3 + N_{34})$ remains constant. As a result, the limiting solubility, polarizations of both phases, and the transition temperature depend on the ratio of the total number of ³He and ⁴He particles in the system, $(N_{34} + N_3)/N_4 = X/(1 - X)$, and therefore on the details of the method of polarization. Of course, in thermodynamic equilibrium this ratio is irrelevant.

Generally speaking, all models which apply at arbitrary polarization give different estimates for the size of the change in solubility, causing a rather large spread in values for T_c . However, all models share the common property that a sizeable increase in polarization results in an increase in T_c by orders of magnitude, whatever the mechanism.

Concluding, we might say that spin polarization is a probable way of lifting the superfluid transition temperature up to experimentally accessible temperatures, although theories are inconclusive about the mechanism which causes the enhancement of T_c . The results lead us to conclude that around 1 MPa one can expect the highest transition temperatures. Also, the overall concentration X in the sample cell might have a large effect on the value of T_c . In spite of the theoretical uncertainties, we feel that investigation of dense polarized ³He-⁴He mixtures is strongly encouraged by the results.

ACKNOWLEDGMENTS

One of us (L. P. R.) would like to thank Prof. R. de Bruyn Ouboter for helpful discussions concerning the properties of the helium isotopes at zero temperature, Prof. F. Pobell for careful reading of the manuscript and the Alexander von Humboldt Stiftung, Bonn, Federal Republic of Germany for awarding him a fellowship which enabled the completion of this paper. Two of the authors (A. E. M. and K. S. B.) would like to thank Prof. G. Frossati and his group for the hospitality during their visits to the Kamerlingh Onnes Laboratory and the Lorentz Institute while this work was being completed. A. E. M. is also grateful to NATO and NSF for support (grants CRG-910283 and DMR-9412769). This work is part of the research programme of the Stichting F.O.M., which is financially supported by the N.W.O., the Netherlands.

REFERENCES

- 1. J. R. Owers-Bradley, H. Chocolacs, R. M. Mueller, C. Buchal, M. Kubota, and F. Pobell, *Phys. Rev. Lett.* **51**, 2120 (1983).
- 2. H. Ishimoto, H. Fukuyama, N. Nishida, Y. Miura, Y. Takano, T. Fukuda, T. Tazaki, and S. Ogawa, J. Low Temp. Phys. 77, 133 (1989).
- G. Oh, M. Nakagawa, H. Akimoto, O. Ishikawa, T. Hata, and T. Kodama, *Physica B* 165-166, 527 (1990).
- 4. R. König and F. Pobell, Phys. Rev. Lett. 71, 2761 (1993).
- 5. G.-H. Oh, Y. Ishimoto, T. Kawae, M. Nakagawa, O. Ishikawa, T. Hata, and T. Kodama, J. Low Temp. Phys. 95, 525 (1994).
- 6. P. Nacher, I. Shinkoda, P. Schleger, and W. Hardy, Phys. Rev. Lett. 67, 839 (1991).
- L. P. Roobol, S. C. Steel, R. Jochemsen, G. Frossati, K. S. Bedell, and A. E. Meyerovich, Europhys. Lett. 17, 219 (1992).
- 8. G. A. Vermeulen, J. Low Temp. Phys. 94, 5 (1994).
- 9. G. Tastevin, J. Low Temp. Phys. 89, 317 (1992).
- 10. D. Candela, M. E. Hayden, and P. J. Nacher, Phys. Rev. Lett. 73, 2587 (1994).
- 11. J. Bardeen, G. Baym, and D. Pines, Phys. Rev. Lett. 17, 372 (1966).
- 12. P. G. van de Haar, G. Frossati, and K. S. Bedell, J. Low Temp. Phys. 77, 35 (1989).
- 13. K. S. Bedell and C. R. Sanchez-Castro, Phys. Rev. Lett. 57, 854 (1986).
- 14. C. R. Sanchez-Castro, K. S. Bedell, and S. A. J. Wiegers, Phys. Rev. B 40, 437453 (1989).
- 15. G. K. Walters and W. M. Fairbank, Phys. Rev. 103, 262 (1956).
- 16. D. O. Edwards, E. Ifft, and R. Sarwinski, Phys. Rev. 177, 380 (1969).
- 17. S. Yorozu, M. Hiroi, H. Fukuyama, H. Akimoto, H. Ishimoto, and S. Ogawa, *Phys. Rev.* B 45, 12942 (1992).
- 18. D. O. Edwards and J. G. Daunt, Phys. Rev. 124, 640 (1961).
- 19. P. Seligmann, D. O. Edwards, R. E. Sarwinski, and J. T. Tough, *Phys. Rev.* 181, 415 (1969).
- D. O. Edwards, D. F. Brewer, P. Seligmann, M. Skertic, and M. Yaqub, *Phys. Rev. Lett.* 15, 773 (1965).
- 21. R. de Bruyn Ouboter and C. N. Yang, Physica B 144, 127 (1987).
- 22. K. Huang and C. N. Yang, Phys. Rev. 105, 767 (1957).
- A. A. Abrikosov and I. M. Khalatnikov, Zh. Eksp. Teor. Fiz. 32, 1083 (1957), [Sov. Phys. JETP 5 (1957) 887].
- 24. J. H. Ager, R. M. Bowley, R. König, and J. R. Owers-Bradley, Phys. Rev. B 50, 13062 (1994).

- 25. E. P. Bashkin and A. E. Meyerovich, Adv. Phys. 30, 1 (1981).
- G. Baym and C. J. Pethick, Landay Fermi-liquid theory (John Wiley and Sons, Inc., New York, 1991), (and references therein).
- A. E. Meyerovich, Spin-Polarized Phases of ³He (North Holland, Amsterdam, 1990), W. P. Halperin and L. P. Pitaevskii, eds. (and references therein).
- 28. A. E. Meyerovich and K. A. Musaelian, Phys. Rev. Lett. 74, 1710 (1994).
- 29. L.-J. Wei, N. Kalenchofsky, and D. Candela, Phys. Rev. Lett. 71, 879 (1993).
- 30. J. Owers-Bradley, A. Child, and R. M. Bowley, Physica B 194-196, 903 (1994).
- 31. H. Ramm, P. Pedroni, J. R. Thompson, and H. Meyer, J. Low Temp. Phys. 2, 539 (1970).
- 32. J. R. Thompson Jr., H. Ramm, J. F. Jarvis, and H. Meyer, J. Low Temp. Phys. 2, 521 (1970).
- A. I. Ahonen, M. A. Paalanen, R. C. Richardson, and Y. Takano, J. Low Temp. Phys. 25, 733 (1976).
- 34. S. A. J. Wiegers, P. E. Wolff, and L. Puech, *Phys. Rev. Lett.* 66, 2895 (1991), [Commented by M. T. Béal-Monod and E. Daniel, *Phys. Rev. Lett.* 68 (1992) 3817].
- 35. B. Castaing and P. Nozières, J. Phys. (Paris) 40, 257 (1979).
- 36. S. A. J. Wiegers, C. C. Kranenburg, T. Hata, R. Jochemsen, and G. Frossati, *Europhys. Lett.* 10, 477 (1989).
- 37. F. Dalfovo and S. Stringari, J. Low Temp. Phys. 71, 311 (1988).
- 38. K. S. Bedell, Phys. Rev. Lett. 54, 1400 (1985).
- 39. K. S. Bedell and P. G. van de Haar, Europhys. Lett. 14, 469 (1991).
- 40. D. S. Greywall, Phys. Rev. B 33, 7520 (1986).
- 41. B. R. Patton and A. Zaringhalam, Physics. Lett. 55A, 95 (1975).
- 42. M. Y. Kagan and A. Chubukov, Sov. Phys. JETP 50, 517 (1989).